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Exo-MerCat v2.0.0: updates and open-source release of the Exoplanet Merged Catalog software

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Abstract

Exoplanet research is at the forefront of contemporary astronomy recommendations. As more and more exoplanets are discovered and vetted, databases and catalogs are built to collect information. Various resources are available to scientists for this purpose, though every one of them has different scopes and notations. In Alei et al. (2020) we described Exo-MerCat, a script that collects information from multiple sources and creates a homogenized table. In this manuscript, we announce the release of the Exo-MerCat v2.0.0 script as an upgraded, tested, documented and open-source software to produce catalogs. The main upgrades on the script concern: 1) the addition of the TESS Input Catalog and the K2 Input Catalog as input sources; 2) the optimization of the main identifier queries; 3) a more complex merging of the entries from the input sources into the final catalog; 4) some quality-of-life improvements such as informative flags, more user-friendly column headers, and log files; 5) the refactoring of the code in modules. We compare the performance of Exo-MerCat v2.0.0 with the previous version and notice a substantial improvement in the completeness of the sample, thanks to the addition of new input sources, and its accuracy, because of the optimization of the script.

Keywords: (Stars): planetary systems, catalogues, Virtual Observatory Tools

1. Introduction

The last few years have seen a huge advancement in the detection and characterization of extrasolar planets. From the discovery of the first exoplanet in 1995 (Mayor and Queloz, 1995), we discovered more than 7000 candidates and 6000 confirmed exoplanets in our galactic neighborhood – and these numbers increase daily. A variety of space- and ground-based telescopes have discovered a large number of planetary candidates, and more facilities will be doing so in the near- and long-term future. As ancillary resources for any researcher and to store data in a homogenized way, catalogs are developed and maintained. Such resources serve both as input target lists for mission scheduling, as well as output storage from past observations.

In 2020, we released the first version of Exo-MerCat (Alei et al., 2020), a Python script that produces a catalog of known and candidate exoplanets by merging data collected from mul-

tle input sources. As described in Alei et al. (2020), merging entries belonging to different source catalogs proved to be a challenging problem because of the presence of aliases (i.e., the same exoplanet listed in one or more catalogs using different nomenclature). For this reason, Exo-MerCat retrieves a common name for the planet target, linking its host star name with the preferred identifier in the most well-known stellar databases. This makes the merging of different input sources possible and it has allowed the creation of a homogenized catalog containing data from the four major input catalogs used in the community: the *Exoplanets Encyclopaedia*² (Schneider et al., 2011), the *NASA Exoplanet Archive*³ (Akeson et al., 2013), the *Open Exoplanet Catalogue*⁴ (Rein, 2012) and the *Exoplanet Orbit Database* catalog⁵ (Wright et al., 2011). Because of the nature of the merge, this catalog is not self-consistent (i.e., all measurements for the same target do not belong to the same reference paper). Rather, it is a collection of

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²<http://exoplanet.eu/>

³<https://exoplanetarchive.ipac.caltech.edu/>

⁴<https://openexoplanetcatalogue.com>

⁵<http://exoplanets.org/>

the most precise estimates for each planetary parameter (mass, minimum mass, radius, period, semi-major axis, eccentricity, inclination), based on the lowest relative error, and the corresponding reference for each preferred estimate. We refer to Alei et al. (2020) for further details on the implementation of Exo-MerCat v1.0.0.

The catalog produced by Exo-MerCat has been used as a source for input target lists for space-based missions, as well as an additional tool for various publications (see Sect. 3.2). In time, we received specific requests to update the code to accommodate the needs of the current users, which we deemed useful and worthy of a software update. Furthermore, the input sources have undergone significant changes (e.g., the migration of the NASA Exoplanet Archive to a new table and notation), which made the ingestion into Exo-MerCat fail at runtime. Finally, to streamline software updates and to allow continuous maintenance of the code, we decided to release the updated version of Exo-MerCat as open-source code on GitHub.

In this paper, we present Exo-MerCat v2.0.0 as a more optimized, robust, and user-friendly version of the Exoplanet Merged Catalog software. In Sect. 2 we describe the updates that were performed on Exo-MerCat. In Sect. 3 we compare the performance of v2.0.0 compared to the previous version, we discuss the value of such a resource for the community, and we describe the open-source philosophy and maintenance of the script going forward. We provide a summary of the work in Section 4.

2. Exo-MerCat v2.0.0: main features

Every version of Exo-MerCat fulfills the following tasks:

1. Ingesting the input sources as tables. Since different input sources come in different formats, the script can flexibly download the tables and convert them in the same .csv format.
2. Homogenizing the input sources to handle the same type of data, in terms of table columns and units. Different input sources have different data available in various formats and are made as similar as possible to ease the future tasks.
3. Selecting a host star identifier for each target in the tables that is the same across the input sources. This is done by querying the host star with stellar catalogs.
4. Merging the entries from the various input sources into a single entry. When multiple estimates for the same planetary parameter are available, Exo-MerCat prefers the most precise measurement.

While this has not changed between the various versions, in Exo-MerCat v2.0.0 these tasks are executed in a more streamlined and optimized way. We summarize the main differences between Exo-MerCat v2.0.0 and its previous version in Table 1. Each process in the itemized list and in the table is described in more detail in Sections 2.1 to 2.5.

The Exo-MerCat script is written in Python and it has been tested for Python v3.8 and above.

2.1. Input sources ingestion

Exo-MerCat v2.0.0 takes into account the following input sources: the NASA Exoplanet Archive (hereafter NASA); the Exoplanet Encyclopaedia (hereafter EU); the Open Exoplanet Catalogue (hereafter OEC); the EPIC/K2 Planets and Candidates Table⁶ (hereafter EPIC); and the TESS Project Candidates Table⁷ (hereafter TOI). Like in Exo-MerCat v1.0.0, the Kepler Objects of Interest Table⁸ (hereafter KOI) is used as a look-up table to check and homogenize the status of the Kepler targets.

In this new version of the software, we kept three core catalogs from v1.0.0 (NASA, EU, OEC). We removed the Exoplanet Orbit Database from the input sources, since the table has been retired, and additional tables were either added from scratch (in the case of the TOI table) or promoted from look-up table to input sources (in the case of EPIC). We chose to add or promote these tables since they are frequently updated and contain the latest updates on the TESS mission results and on further analyses on K2 data. The Kepler Objects of Interest table has not been updated since 2018, as all vetting has been completed. However, the KOI table is still useful as a look-up table since it reports all the updated statuses of each target in the Kepler sample and it is able to fix discrepancies in the other catalogs.

Some of the input tables changed between Exo-MerCat v1.0.0 and the current version. The NASA Exoplanet Archive retired the tables that were used by Exo-MerCat v1.0.0 for the NASA, KOI, and EPIC; new tables are hosted now on different URLs. Some column names were replaced and other columns were removed/joined together. Furthermore, the NASA catalog (the `pscompars` table) now contains heterogeneous data from different sources. For more details, we refer the reader to the NASA Exoplanet Archive website⁹. Similarly, the EU catalog has undergone major updates in their database structure and layout; some column names were changed and the relative URL was modified. These updates required an adjustment of the input source ingestion to ensure that the Exo-MerCat script would not fail at runtime. The changes consisted in updating the URLs and the naming conventions according to the documentation provided by the NASA Exoplanet Archive and the Exoplanet Encyclopaedia.

Finally, the Open Exoplanet Catalogue now provides a cumulative .xml file in addition to the separate .xml files for each system. Exo-MerCat has been updated to download the cumulative file instead.

In Exo-MerCat v1.0.0, the URLs that allowed the download of the input source tables were hard-coded. In v2.0.0, Exo-MerCat reads the `input_sources.ini` file to get this information, as well as the destination and name of the downloaded file. The user can update the links to the source files by

⁶<https://exoplanetarchive.ipac.caltech.edu/cgi-bin/TblView/nph-tblView?app=ExoTbls&config=k2pandc>

⁷<https://exoplanetarchive.ipac.caltech.edu/cgi-bin/TblView/nph-tblView?app=ExoTbls&config=TOI>

⁸<https://exoplanetarchive.ipac.caltech.edu/cgi-bin/TblView/nph-tblView?app=ExoTbls&config=cumulative>

⁹<https://exoplanetarchive.ipac.caltech.edu/docs/transition.html>

Exo-MerCat v1.0.0	Exo-MerCat v2.0.0
Input sources ingestion (Sect. 2.1)	
Hard-coded links to sources	User-specified links to sources (to be configured in <code>input_sources.ini</code>)
Deprecated links to input sources	Updated input sources download
Exoplanet Orbit Database as source	Removed Exoplanet Orbit Database from input sources
–	Option to load specific versions of the catalogs when available in the <code>InputSources/</code> folder
Input sources homogenization (Sect. 2.2)	
Default .0x to letter replacements (.01=b,.02=c...)	<i>*performed later</i>
Potential Brown Dwarf: set letter value as “b”	Potential Brown Dwarf: set letter value as “BD”
Hard-coded replacements	User-specified replacements
KOI look-up table: replace updated status	KOI look-up table: save both old and updated status
Host star identifier selection (Sect. 2.3)	
Check that coordinates match within 36 arcsec; replace mismatches with mode or mean	Check that coordinates match within 1 arcsec; log disagreements, but no replacement applied
Main Identifier: Query by host in SIMBAD. If not found, query by coordinates at increasing radii (up to ≈ 10 arcsec) in SIMBAD, KIC, K2, Gaia until all are found	Main Identifier: Query by host in SIMBAD and TIC. If not found, query by coordinates at increasing radii (up to 1 arcsec) on SIMBAD and TIC. If not found, use catalog identifier
Check duplicated entries	<i>*performed later</i>
Merging the entries (Sect. 2.4)	
<i>*performed earlier</i>	Check for inconsistent <code>letter</code> values by grouping by (<code>main_id</code> , <code>binary</code> and <code>p</code> or <code>a</code>); replace “.0d” into letter when possible and check inconsistent letters
Group entries on (<code>main_id</code> , <code>binary</code> and <code>letter</code>) for merging	Group entries on (<code>main_id</code> , <code>binary</code> , <code>letter</code>) for merging, then divide in subgroups by <code>p</code> or <code>a</code>
<i>*performed earlier</i>	Check for duplicated values in merging entries
If status is not consistent, prefer FALSE POSITIVE or CANDIDATE if at least one is present	If status is not consistent, write CONTROVERSIAL.
Alphanumeric string to represent status of merged entries	User-friendly status string
Catalog name selected from available aliases	Common naming convention: <code>main_id</code> + <code>binary</code> + <code>letter</code>
–	Use of flags to highlight errors in the merging process for easy filtering
Outputs (Sect. 2.5)	
No brown dwarf filtering	Possible filtering of brown dwarfs based on mass criterion (m or $m \sin i < 20M_{Jup}$)
Logging provided only at run-time	Logging provided at run-time (if desired), plus saved in ancillary files for asynchronous reference

Table 1: Summary of the main differences between Exo-MerCat v1.0.0 and Exo-MerCat v2.0.0. The star and bullet symbols link the tasks that are performed at different points in the script in the two versions.

editing the `input_sources.ini` file. In this way, it will be possible to quickly adapt to changes in URLs and addition/removal of specific columns, or input catalogs directly.

By default, the catalogs are downloaded directly into the `InputSources` folder. The file name includes the download date in `YYYY-MM-DD` (month, day, year) format. When running the script, the user can specify the flag `-date YYYY-MM-DD` (or `--d YYYY-MM-DD`) to load the version of the input catalogs downloaded at a specific date, provided they are present in the `InputSources` folder. If the download of a current version of any of the six tables fails at runtime, the run will continue provided that previous tables are present in the folder. In this case, the script automatically prints a warning for the user and loads the most recent table available. If no earlier tables are available, the script fails at runtime. Issues on the download of the input catalogs are often temporary and will likely be fixed at the source in a short amount of time.

2.2. Input sources homogenization

Each input source collects specific datasets for a selected amount of targets. The choice of including or excluding a specific target depends on the philosophy of each catalog: for example, the NASA catalog only displays confirmed planets; the EU catalog collects also candidate and free-floating planets, as well as Solar System objects and trojans. Various catalogs could also adopt different mass cutoffs when it comes to the distinction between exoplanets and brown dwarfs (see Alei et al., 2020, for details). We cannot expect to have the same information among different catalogs. Still, the format of each table can be homogenized to allow a fairer comparison and merging of the entries that are included in two or more catalogs.

The Exo-MerCat script performs specific operations for each input catalog to ensure that the homogenized output can easily allow for a comparison among the various tables. First, it selects and renames all the columns in the original files that will be used for the merging. These include: the planet name and its host star; the aliases by which the planet and/or the star are known; the planet letter and the binary string associated to the host star; the discovery method and year; and the estimates (including uncertainties and reference) for planetary mass, minimum mass, planetary radius, semi-major axis, eccentricity, period, and inclination.

We do not include stellar parameters or radial velocity and transit observables since those are not uniformly stored in all input sources, or rely on self-consistent measurements that cannot be provided in this joint catalog per definition. For these parameters, we recommend querying specific stellar or planetary catalogs.

After ingestion, the planet and star names are homogenized to have a similar notation in the various catalogs. Normally, a confirmed exoplanet is known by the name of its host star followed by a lowercase letter (most commonly, “b” for single-planet systems). When there are multiple planets orbiting the same host star, subsequent letters from “b” to “j”¹⁰ are used to

label all the planets in the system, by year of discovery. Candidate exoplanets, generally revealed by transit detection space missions, are instead labeled as the name of their host star in the input catalog of the mission followed by the string “.0d” where “d” represents an integer that starts from 1 and increases as more candidates are found in the same planetary system. These candidates, once vetted by complementary methods, are then confirmed or rejected. In case of confirmation, they then become known with their planetary letter.

In Exo-MerCat v1.0.0 the candidate “.0d” indexing was replaced by the corresponding letter (assuming that “.01” becomes “b”, “.02” becomes “c”...). However, in reality this relationship is not so obvious, as the vetting process might take more or less time depending on the candidate planet, and automatically transforming the candidate letter into the confirmed letter might have lead to mismatches in the merges of a few candidates. It might have also been a source of confusion for the user, since candidate planets are not normally found labeled with a letter, but only with “.0d” strings. In the new version of the code, we removed this automatic replacement between numeric string and letter while homogenizing the catalogs to avoid unintentional mismatches. This choice leads to a subsequent issue: the same confirmed planet might have been present in different catalogs with its candidate name (ending in “.0d”) as well as its official name (with the corresponding letter); we resolve these differences when collapsing the duplicated catalog entries into the merged Exo-MerCat entry (see Sect. 2.4).

In Exo-MerCat v1.0.0, we also automatically added the letter “b” to all entries that did not show a lowercase letter nor a “.0d” string in the name. This is often the case for substellar objects. In Exo-MerCat v2.0.0, we reserve the specific value “BD” of the planet letter column (called `letter`) for all entries that do not show a letter nor a “.0d” string in their name. It is thus possible to filter these candidates out easily from the final catalog.

In the case of binaries, the information content in the input sources is not uniform. Specifically, OEC provides a flag that establishes if the planet orbits in a P-type (circumbinary) or S-type orbit (around one of the binary companions), but it is not often clear about which companion hosts the planet. In Exo-MerCat v1.0.0, the `binary` value that was only identified as “S-type” was discarded from the homogenized output. In Exo-MerCat v2.0.0, this partial information is still retained and, when possible, corrected at a later stage (see Sect. 2.3).

As was done in v1.0.0, the Exo-MerCat script checks that the status associated with the entries in the input catalogs is up-to-date. This can easily be done for the Kepler targets (the majority of the confirmed planets known so far) by comparing the value of the status of each Kepler target included in the input sources with the status in the KOI table. This is relevant especially in the case of candidates that were then revealed as false positives. While the status might not have been updated in some input sources we consider in Exo-MerCat, it has been updated in the KOI table, so we refer to that as the most accurate status. As of v2.0.0, both the original status and the updated status after the check on the KOI table is stored in Exo-MerCat in the `original_status_string` and `checked_status_string`

¹⁰As of October 25, 2024, there is only one occurrence of a letter “j”, for the candidate planet HD 10180 j.

columns.

In Exo-MerCat v1.0.0 we implemented hard-coded modifications to some specific targets in the sample that appeared to possess errors in their input catalogs, or to manually homogenize names of special targets (e.g., the planets in the PSR 1257+12 system, known with the suffixes “A”, “B”, and “C”, normally used for stellar binary companions). These corrections were performed at various points within the script. This put an additional layer of complexity for any user wanting to customize such replacements. As of v2.0.0, any user can force the replacement of a specific cell value (name, host star, binary letter, or coordinates) by providing the `replacements.ini` file. This action might be necessary to: 1) replace known mistakes on names of planets or host stars that have been known to occur in the input catalogs; 2) drop selected elements from the list (e.g., Solar System items, which are also included in the EU catalog); 3) force a specific value of the binary suffix (see Sect. 2.2); 4) manually set specific right ascension and declination values. The `replacements.ini` file will surely change with version updates, as the input sources evolve. Users are invited to modify and submit any additional replacement to be performed via push request (see Sect. 3.3).

The log files that are produced at run time provide information on any additional replacement to be performed or if any replacement is no longer necessary. In the latter case, the `replace_known_mistakes.txt` log file lists all the replacements that were not used when homogenizing each source. We expect to always have some entries in this file, since some replacements are specifically tailored for one source catalog and might not be necessary for all the others. However, the user could optimize the contents of `replacements.ini` if no replacement is performed in any source. Some replacements could also no longer be necessary if the issue was fixed by the source catalog at some point in time.

In all versions of Exo-MerCat, the coordinates are converted to decimal degrees for KOI and OEC. Furthermore, theoretical estimates of masses and radii are removed. In Exo-MerCat v1.0.0 all parameter estimates that had infinite or null errors were also removed. This is no longer the case for Exo-MerCat v2.0.0, which retains partial information when available.

Separate files for the homogenized entries for each input catalog are then collected in the `StandardizedSources` folder as `.csv` files.

2.3. Host star identifier selection

In every version of Exo-MerCat, once the input catalogs have been homogenized, they are concatenated in a joint pandas DataFrame. In principle, the same target can appear in the concatenated DataFrame up to five times, depending on how many input sources contained it. Since the same planet can be labeled differently depending on naming conventions and host star aliases, Exo-MerCat performs various operations to assign a single identifier to each host star, to ensure a successful merging of each entry (see Sect. 2.4).

All versions of Exo-MerCat perform a check to identify matches between the `host` column and the known aliases

provided by each input catalog. In Exo-MerCat v2.0.0, the `alias_as_host.txt` log file lists all of the host stars found by another name within the various input catalogs.

In Exo-MerCat v1.0.0, all entries that shared the same host star were checked for coordinate mismatches. The DataFrame was grouped by the host star name; any coordinate that differed more than 0.01 degrees for the same host star was flagged and the coordinates of all entries of the group were changed to the mode value of the disagreeing coordinate. In Exo-MerCat v2.0.0 we perform this test only on smaller fractions of the sample, such as planets orbiting binaries or planets whose main identifier was not found through name search (see below). In every coordinate check, the tolerance is by default smaller (1 arcsec) and the script does not perform any change but simply prints the name of the system in a log file for user vetting (see Sect. 2.5).

Binaries are handled in a slightly different way in Exo-MerCat v2.0.0. Previously, duplicated entries were grouped by host star name and more specific values of `binary` were replaced when more information was available (though excluding “S-type” binaries, see Sect. 2.2). Exo-MerCat v2.0.0 groups by host star name and letter and allows the treatment of “S-type” binaries: “S-type” and null `binary` values are replaced when additional information is present; if only null or “S-type” strings are present, the “S-type” value is preferred since it provides at least partial information.

Two flags are now introduced to highlight the entries that failed binary-related tests. A first coordinate check is performed for all binary systems (see above). Entries that fail the coordinate check will have the value of the `binary_coordinate_mismatch_flag` column set to 1. If there are complex binary systems (e.g., coexistence of “A”, “B”, and/or “AB” binary values for the same `host+letter` group), the value of the `binary_complex_system_flag` column is assigned to be equal to 1. Data is not modified in either of the two tests. The flags are only used to help the user filter out problematic entries at a later stage.

The `check_binary_mismatch.txt` log file lists all the corrections that were done to infer the value of the `binary` column, as well as the results of the tests performed. The user can review the log file and force the value of `binary` in `replacements.ini` if deemed necessary.

The next task is to query stellar catalogs to identify a common main identifier for each target star. To select a main identifier for the host star and the corresponding coordinates, Exo-MerCat v1.0.0 first queried SIMBAD¹¹, then VizieR¹² as a gateway to the Kepler Input Catalog (Kepler Mission Team, 2009), the K2/EPIC Input Catalog (Huber et al., 2017), and Gaia DR2 (Gaia Collaboration et al., 2016, 2018). Exo-MerCat v1.0.0 queried these catalogs by the host star name of each entry; if the query was unsuccessful, it searched again by coordinates, using a “ConeSearch”. The ConeSearch is a specific query that can be performed that can be performed by all Virtual Observatory aware services, which consists in searching by

¹¹<https://simbad.cds.unistra.fr/simbad/>

¹²<http://vizier.u-strasbg.fr/viz-bin/VizieR>

coordinates for the closest star inside a cone of a given tolerance radius. In Exo-MerCat v1.0.0 the radius of the cone was gradually increased until all identifiers were found. This meant that, for the entries that did not have any counterpart available in the stellar catalogs, the script would associate a star that could have been far away from the planet candidate. Although this was the case for less than 100 identifiers in the cumulative catalog, it led to an erroneous result.

In Exo-MerCat v2.0.0, we only consider SIMBAD and TIC v8.2 (Paegert et al., 2021) as our stellar sources. Exo-MerCat v2.0.0 first queries SIMBAD for the host name and the available aliases for each entry. The query also collects the official coordinates (epoch J2000) in degrees from SIMBAD. This query is performed using all the available identifiers for a given target until a match is found, and including any information about binary stars if available. For all entries for which a match is found in SIMBAD searching by identifiers, we associate a value of `angular_separation` equal to 0.0 (since no coordinates were involved in the query) and we specify SIMBAD in the `main_id_provenance` column. The same query is repeated searching on the TESS Input Catalog (v8.2) for all entries whose SIMBAD search was unsuccessful. This catalog is the most recent input catalog available and it logs all the GAIA DR2 identifiers and coordinates, in addition to the TIC identifiers (TICid). If a match is found in TIC searching by TICid for a given entry, `angular_separation` is set to 0.0 and `main_id_provenance` is set to TIC.

For the targets for which no SIMBAD nor TIC identifiers have been found, the script prepares for the query on the stellar catalogs based on coordinates. First, a coordinate check is performed (see above). If either the right ascension, the declination, or both coordinates do not match for all the entries reported to be orbiting the same host star, the flag `coordinate_mismatch_flag` will be set to 1, while the column `coordinate_mismatch` will store RA, DEC, or RADEC to inform on the specific coordinate(s) that failed the check. Entries that failed this check are also printed in the `check_coordinates.txt` file for user reference. No data is modified during this check.

A 1 arcsec radius ConeSearch on SIMBAD is then performed. The main identifier is selected to be the one with the smallest angular separation among all sources within the ConeSearch. The separation values for all successful queries are saved in the `angular_separation` column and `main_id_provenance` is set to SIMBADCOORD. Then, the same ConeSearch query is repeated on the TIC catalog for all the targets still missing a main identifier. Also in this case, if the 1 arcsec ConeSearch is performed successfully, the `angular_separation` column is updated and the value of the column `main_id_provenance` is set to TICCOORD. When identifiers are not found either through a name search or a ConeSearch in SIMBAD or TIC assuming a radius of 1 arcsec, Exo-MerCat v2.0.0 does not increase the radius any further but assumes as main identifier and coordinates the ones stored in the input catalogs. In this case the `angular_separation` column is set to -1 and the `main_id_provenance` is filled with a string labeling the input catalog from which the identifier is

taken.

In Exo-MerCat v2.0.0, once the main identifier has been found, the script cleans the `main_id` values to ensure that no exoplanet identifiers (a small fraction of them also included in SIMBAD) are considered. The script also collects any extra information that the main identifier can provide, adding details on the binarity of the host star in the `binary` column. This is automatically done if the current value of `binary` is “S-type” or null. If the value in `binary` is non-null and disagrees with what found by the query on the stellar catalogs, this discrepancy is logged in the `polish_main_id.txt` file, recommending user supervision. Then, the checks on the binary stars are repeated using the main identifier as a criterion for the grouping of the cumulative catalog (i.e. by grouping on `main_id + letter`), to identify any further potential binary coordinate mismatches or complex systems.

Further checks are then performed to: 1) ensure that the same host star name is not found with different main identifiers, which would impair a correct merging of the entries (see Sect. 2.4); 2) ensure that the the same identifier is associated with the same coordinates, and 3) ensure that the same aliases are collected for the same main identifier. While these tests do not modify the catalog, they can be a useful diagnostic to the user in case issues occur in the stellar catalog query, or to identify peculiar cases. The results of these tests are logged in the `post_main_id_query_checks.txt` file.

2.4. Merging

What makes Exo-MerCat useful is the merging of heterogeneous information from different source catalogs. For this reason, we have developed the merging algorithm that was used ever since the first release of the code. The goal of the merging is to obtain one single entry per target that collects the best estimate for each relevant parameter (mass, minimum mass, radius, period, semi-major axis, inclination, eccentricity) from the source catalogs and their references. The best estimate is the one that has the lowest relative error (see Alei et al., 2020, for details). In v2.0.0, we optimized this module to account for all the updates we implemented and to make the merging process more robust, by relying on physical parameters in addition to the (homogenized) names of each target.

Since we handle the exoplanet letter differently in Exo-MerCat v2.0.0 (see Sect. 2.2), the same target could be present in the concatenated DataFrame with different values `letter` (e.g. with a “.0d” string as well as with a lowercase letter). For this reason, Exo-MerCat v2.0.0 first checks that the planet letter is the same for every entry having the same main identifier, binary string, and planetary period (or semi-major axis when no available period measurement is available). If only one lowercase letter appears for each group, any “.0d” string for the same group is replaced with the same lowercase letter. If there are more than one lowercase letter values that are inconsistent, these are not changed in the catalog, but are logged in the `group_by_period_check_letter.txt` file and should be vetted by the user.

In Exo-MerCat v1.0.0, the merging was executed based on the `main_id`, the `binary`, and the `letter` columns.

Exo-MerCat v2.0.0 adds another layer of complexity to the merging to ensure that no mismatches are accidentally performed. The newly implemented merging routine checks that each group of entries that share `main_id`, `binary`, and `letter` also have the same period. If this is the case, the merging is performed. If the period values are different more than 10% of the value of each estimate (though the tolerance can be changed by the user), the merging is performed on the subgroups that share the same period value. In the case where the period is not available, the grouping is performed using the semi-major axis. In the case where neither the period nor the semi-major axis are available, the merge is forced over all the elements of the group (hereafter: “fallback merge”).

All occurrences where there is a disagreement on period and/or semi-major axis, or when a fallback merge occurs, are logged in the `group_by_letter_check_period.txt` file. These occurrences are also recorded in two flags: the `period_mismatch_flag` column, which is set to 1 when there has been a disagreement for the period or semi-major axis and multiple entries have been created; and the `fallback_merge_flag`, set to 1 when no information on period or semi-major axis was available and the merge was forced.

The merging module itself has undergone some minor updates. First of all, the naming convention changed: while in Exo-MerCat v1.0.0 a name was picked as the most common one from the input catalogs, Exo-MerCat now creates the preferred target name by joining the main identifier, the binary value, and the letter value. The names by which every target was known in the original catalogs are also stored in their respective columns (see Table B.4), a new addition to v2.0.0. Users can therefore query the input catalogs by name to gather further information or to retrieve the original dataset.

The determination of the status was also updated, as two additional status flags (“controversial” and “preliminary”) are now available. Exo-MerCat sets “controversial” status label for every entry that did not show agreement in the `checked_catalog_status` (i.e., the updated status after the check on the KOI table). The label “preliminary” is instead reserved to elements in the source catalogs that have a null status. These are entries in the source tables that have been very recently added and are supposed to be updated in the near future. The original and checked catalog status values per each catalog are also collected in the `original_status_string` and `checked_status_string` as a string representation of a dictionary (e.g., “eu: confirmed, oec: confirmed”). This improves the readability and user-friendliness of the output catalog compared to Exo-MerCat v1.0.0, where the status was represented as a more cryptic AXDXEXCX string (“A”, “D”, “E”, “C” representing the four source catalogs and “X” being a number between 0 and 2 to represent the status of the planet in each source, see Alei et al., 2020, for details).

The `discovery_method` column (whose header changed for readability, see Table B.4) now allows for multiple detection methods to be listed, in case of independent detections.

When creating the single entry, the script checks that there are no duplicate values within the same input source. In case this happens, these are stored in

the log file `merge_into_single_entry.txt`, the flag `duplicate_catalog_flag` is set to 1 and the duplicate names are stored in the `duplicate_names` column for easy filtering. This is mostly the case when both the candidate and the official planet name are stored in the catalogs. Finally, the output catalog is checked to identify potential duplicate planets that have been not identified during the merging. This happens when there is missing information or discrepancies in the labeling of the binary star or the planet letter. The output catalog is grouped by `main_id`, then periods and semi-major axes are checked. If, in each group, similar values of `p` or `a` are found for entries that do not have agreeing values of `binary` or `letter`, the flag `misnamed_duplicates_flag` is set to 1. The output of this test is logged in `identify_misnamed_duplicates.txt`. These cases should be vetted by the user and replacements should be listed in the `replacements.ini` file to facilitate a correct merge.

2.5. Outputs

In Exo-MerCat v1.0.0, we did not implement any cutoff on the mass and we inherently assumed the heterogeneous cutoffs of the input catalogs. In v2.0.0, we separate the brown dwarf candidates from the merged catalog based on a mass cutoff criterion of $20 M_{Jup}$.

A single run of the Exo-MerCat v2.0.0 script produces three main files: first, the `exo-mercat_full.csv` file contains all the entries that Exo-MerCat produced. Then, two ancillary files are produced: the `exo-mercat_brown_dwarfs.csv` file contains all the entries whose (minimum) mass exceeded $20 M_{Jup}$; the `exo-mercat.csv` file contains all the other entries, including those that Exo-MerCat identified as potential brown dwarfs (`letter` = “BD”) but whose (minimum) mass was below the $20 M_{Jup}$ cutoff. The user can filter the potential brown dwarfs out of the `exo-mercat.csv` file by filtering the value of the `letter` column *a posteriori*, or by forcing the removal of a specific target through the use of the `replacements.ini` file (see Sect. 2.2). By definition, the sum of the two ancillary files returns the full catalog.

When an earlier output catalog file produced by Exo-MerCat is available, the script also compares row by row the current output catalog with the most recent available version and updates the value of the `row_update` column to the current date if any value in the row changed. This allows to keep track of the changes to every single row and to isolate the updated rows in every run.

The names of the column in the catalog generated by Exo-MerCat were updated in v2.0.0 to be more intuitive, and additional columns were added to the table. We explain the translation between the two versions in Appendix A (Table B.4). As stated in the previous sections, the Exo-MerCat v2.0.0 script produces ten log files which inform the user whenever relevant changes on the data have been performed by the Exo-MerCat script, or whenever some flags have been raised. The names of the log files follow the ones of the functions that produced them. In Table 2 we summarize name and function of each log file. In addition to the log files, Exo-MerCat v2.0.0

allows for on-screen logging using the `logging` package. This type of logging can be enabled when running the script with a command-line option `--verbose` or `-v`.

2.6. Folder structure

In this new version, the user can impact the performance of `Exo-MerCat` according to the specific needs. It is also easy to add custom sources to the calculation, if desired.

The script was refactored into modules and classes, included in the main module folder `exomercat`. Documentation and typesetting hints were also added, to increase readability. The six input catalogs, as well as the output catalog generated by `Exo-MerCat`, are now subclasses of a parent class named `Catalog`. The `Catalog` class contains all the common methods for the currently implemented catalogs. These will be read when homogenizing the specific input source unless overwritten by one other subclass, specific to that input source. The subclasses (one per catalog and in different python files) describe all the catalog-specific methods. These methods will raise a `NotImplementedError` if not implemented in the subclass. This way of refactoring the script paves the way for more input sources to be added in the future. The `NotImplementedError` can be used to point out all the methods that need to be specified when developing the ingestion of a new catalog source.

The utility functions are collected as static methods within the `UtilityFunctions` class. These are service functions that are used by one or more classes to perform specific tasks (e.g. read a `.ini` file, gather information from `.xml` files).

The command-line functions that can be used to run `Exo-MerCat` are stored in the `cli.py` file. These include: 1) the `ping` function, used to run checks on the current status of the input sources and the stellar catalogs; 2) the `input` function, which downloads and homogenizes the tables (see Sects. 2.1 and 2.2); the `run` function, which loads homogenized tables and creates the output catalog (see Sects. 2.3 to 2.5); 4) the `check` function, which runs general sanity checks on the output catalog.

During runtime, the most up-to-date versions of the source catalogs are downloaded and stored in the `InputSources` folder. The source catalogs that were successfully refined to a common output layout are stored in the `StandardizedSources` folder. For each run, a set of log files is produced in the `Logs` folder. At the end of every run, the catalog outputs are stored in the `Exo-MerCat` folder. When saving, the input sources and the merged catalog file names contain also the date of the script execution in `YYYY-MM-DD` format.

The released version has been thoroughly tested with unit tests, which are stored in the `tests` folder.

The `.ini` configuration files used to configure the download of the input sources (`input_sources.ini`) and to force known replacements to be performed (`replacements.ini`) are by default in the root folder, together with a `README` file that lists the major changes in each version. If `Exo-MerCat` is installed as a package, these files could be moved to any location¹³.

3. Discussion

3.1. Comparison to previous version

While the philosophy behind `Exo-MerCat` has not changed, the `Exo-MerCat v2.0.0` script produces a catalog that contains more data (because of the inclusion of more input sources) and a more complete set of measurements for each target, compared to the previous versions. In the following, we compare the catalog generated by `Exo-MerCat v2.0.0` with the one generated by the previous version of the script, which has been updated in order to: 1) exclude the retired `Exoplanet Orbit Database` catalog; and 2) replace the URLs to the `NASA`, `EU`, and `KOI` catalogs to make the script not fail at runtime. We refer to this specific older version as `Exo-MerCat v1.1.0`. The catalog generated by `Exo-MerCat v2.0.0` that we use for this comparison is contained in `exo-mercat_full.csv` and it includes also the brown dwarfs, for a fairer comparison with the catalog generated by `Exo-MerCat v1.1.0`. The `Exo-MerCat v2.0.0` and `Exo-MerCat v1.1.0` scripts were run on the same date (October 25, 2024), hence using the same input sources tables (though not all tables were used by `Exo-MerCat v1.0.0`, see Sect. 2.1). Details on the `Exo-MerCat v2.0.0` run as of October 25, 2024 can be found in Appendix B.

In Table 3 we compare the five input catalogs with the ones generated by `Exo-MerCat v1.1.0` and `Exo-MerCat v2.0.0`. The catalog produced with the latest version shows an increase in the amount of planets compared to `Exo-MerCat v1.1.0` of nearly a factor two, mostly thanks to the addition of the `TOI` input source. The number of confirmed planets increases of about 250 targets, while the number of candidate and false positive targets increases of approximately a factor three or more than ten respectively. The entries that have an estimate of the radius or planetary period is also doubled, while the ones that possess a mass or minimum mass estimate are, in number, very similar for the two versions. Again, this variation is connected to the addition of the `TOI` catalog, which provides periods and radii for most of the targets that it contains.

In Figure 1 we compare the distribution of the main planetary parameters between the output catalogs produced by the two versions of `Exo-MerCat`. In this visualization, it is clearly visible the added benefit that `Exo-MerCat v2.0.0` (red histogram) has on the completeness of the sample, especially with respect to the period, semi-major axis, and radius. In Figure 2 we compare the distribution on sky of the catalogs generated by the two `Exo-MerCat` scripts (`Exo-MerCat v1.1.0` in blue, `Exo-MerCat v2.0.0` in red). Here, the larger number of targets included in the catalog generated by `Exo-MerCat v2.0.0` and their position on sky is immediately noticeable and follows the distribution of the `TOI` candidates known to date.

In Figure 3 we show a stacked bar histogram that highlights the distribution of entries in the catalog generated by `Exo-MerCat v2.0.0` depending on the number of input catalogs where these were found (e.g., merged entries that were found in all the input catalogs have a value of 5; entries that were found in only one input catalog have a value of 1). In different colors, we highlight the contribution of the `TOI` and `EPIC` catalogs in the final distribution. In orange, we show all the entries that did

¹³For details, see: <https://exo-mercat.readthedocs.io/>

Log File	Content Summary	Sect.
<code>replace_known_mistakes.txt</code>	Lists unused replacements during the homogenization of the sources.	2.2
<code>alias_as_host.txt</code>	Lists host stars found by other names in input catalogs.	2.3
<code>check_binary_mismatch.txt</code>	Lists corrections for the <code>binary</code> column and warnings for coordinate discrepancies.	2.3
<code>check_coordinates.txt</code>	Contains entries that failed coordinate checks before SIMBAD and TIC queries.	2.3
<code>polish_main_id.txt</code>	Lists entries whose <code>main_id</code> was cleaned and where new <code>binary</code> values were identified.	2.3
<code>post_main_id_query_checks.txt</code>	Lists checks performed after finding main identifier, including duplicate hosts and similar coordinates.	2.3
<code>group_by_period_check_letter.txt</code>	Logs systems with inconsistencies in planet letter for the same period.	2.4
<code>group_by_letter_check_period.txt</code>	Contains groups with disagreements in period or semi-major axis.	2.4
<code>merge_into_single_entry.txt</code>	Lists duplicate entries within the same input catalog.	2.4
<code>identify_misnamed_duplicates.txt</code>	Lists potential missed duplicates in Exo-MerCat because of mismatches in the planet name (<code>binary</code> and <code>letter</code>).	2.4

Table 2: Summary of the log files produced Exo-MerCat v2.0.0 and their function.

Query	NASA	EU	OEC	TOI	EPIC	v1.1.0	v2.0.0
All planets	5785	9806	5402	7241	1800	9900	17445
Confirmed	5785	7373	5288	1012	575	7383	7621
Candidates	0	2329	100	4621	975	2402	7632
False Positives	0	104	14	1176	250	115	1488
With semi-major axis	5499	7083	2802	0	390	6113	8687
With radius	4379	6551	4155	6771	1434	6553	13435
With mass	1999	4811	2505	0	130	4779	5449
With $m \sin(i)$	852	1632	261	0	14	1671	1717
With period	5504	8222	5095	7141	1777	8317	15702
With mass or $m \sin(i)$	2851	4811	2766	0	138	4895	5489
With mass and $m \sin(i)$	0	1632	0	0	6	1555	1677
With mass and $m \sin(i)$ and radius	0	16	0	0	0	18	0
With mass or $m \sin(i)$ and radius	1469	1798	1544	0	127	1847	2487
With mass or $m \sin(i)$, and radius and period	1451	1461	1469	0	127	1519	2105
All systems	4319	7717	4059	6963	1560	7614	14483

Table 3: Available measurements for various combinations of parameters in the five catalogs as they were downloaded from their sources, plus the catalogs produced by Exo-MerCat v1.1.0 and Exo-MerCat v2.0.0. Dataset updated on October 25, 2024.

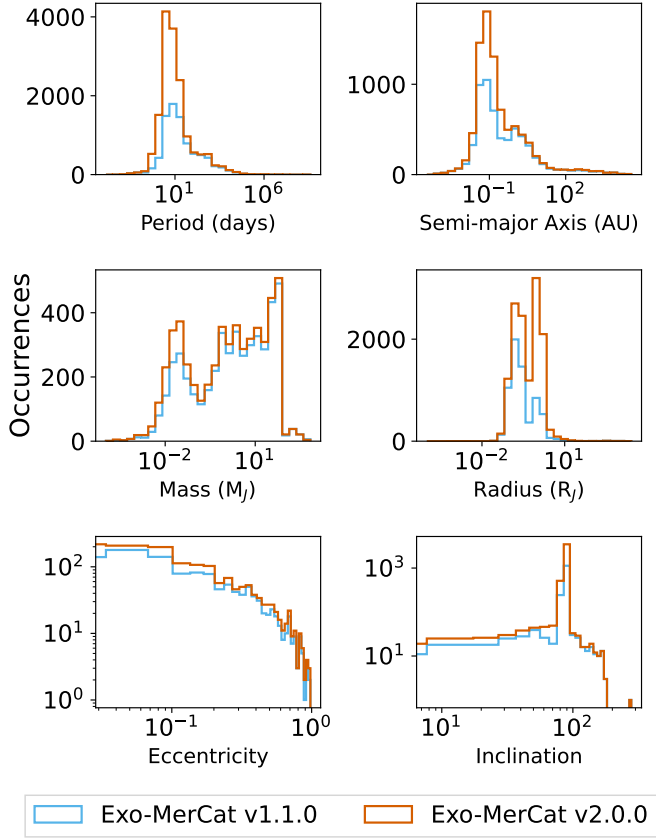


Figure 1: Histogram distributions for the period (top left), semi-major axis (top right), mass (center left), radius (center right), eccentricity (bottom left), and inclination (bottom right) for the catalogs generated by Exo-MerCat v1.1.0 (in blue) and Exo-MerCat v2.0.0 (in red). Dataset updated on October 25, 2024.

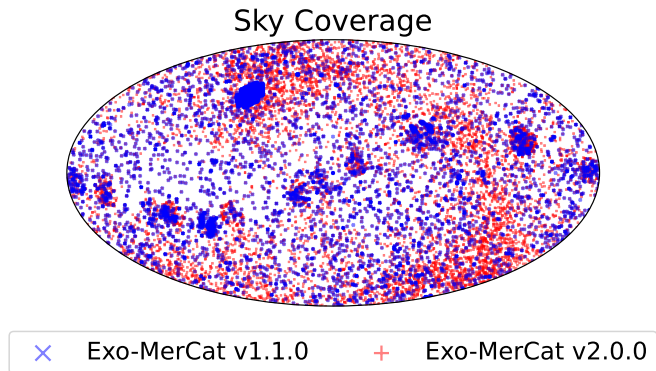


Figure 2: Sky Coverage comparison between the catalogs generated by Exo-MerCat v1.1.0 (in blue) and Exo-MerCat v2.0.0 (in red). Dataset updated on October 25, 2024.

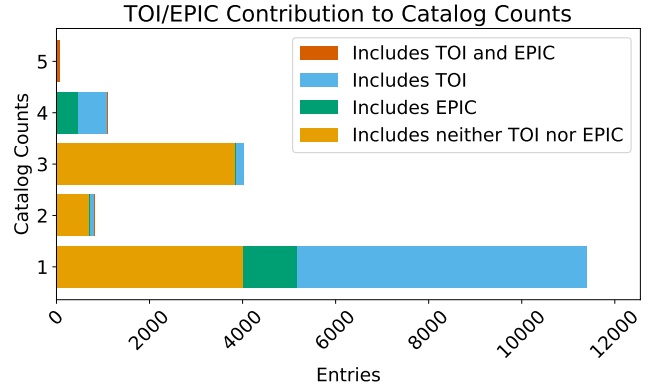


Figure 3: TOI and EPIC contribution to the total sample, measured in catalog counts (from one to five occurrences total in the source catalogs). In red, the sample that includes both TOI and EPIC (potentially together with other input sources); in blue, the sample that includes TOI (potentially together with other input sources) but not EPIC; in green, the sample that includes EPIC (potentially together with other input sources) but not TOI; in orange, the sample that does not include TOI nor EPIC. Dataset updated on October 25, 2024.

not have a counterpart in TOI nor EPIC, but were only found in one or more of the NASA, EU, and OEC catalogs. Because of this, we can only see entries in the orange subset that have a value smaller or equal to 3. In blue, we highlighted the targets that were found in TOI but not in EPIC, while we represent the targets found in EPIC but not in TOI in green. The targets in the green and blue sample can be found also in NASA, EU, and OEC (catalog count is 4), though the majority are found only in TOI or EPIC respectively (catalog count is 1). A small portion of targets is found in both EPIC and TOI, shown in red in Figure 3. This subset is divided into 80 targets that have been found in all catalogs (catalog count is 5) and 38 that have been found only in TOI and EPIC (catalog count is 2). We observe that the contribution of TOI and EPIC catalog does not overlap substantially with the other catalogs, since most of the entries appear in the single-catalog bar (catalog count is 1). However, there is a fraction of candidates that are found in NASA, EU, and OEC, as well as one or both TOI and EPIC (catalog counts is 4 or 5). In this case, the information included in the two additional catalogs has enriched the content of the existing entries, which is one of the fundamental differences between Exo-MerCat v1.1.0 and Exo-MerCat v2.0.0.

3.2. Value to the community

Since the release of v1.0.0, Exo-MerCat has been used in a few publications. For example, Desidera et al. (2023), Benatti et al. (2021), and Carleo et al. (2021) used the catalog generated by Exo-MerCat to identify a subset of known transiting planets that satisfied specific conditions on mass, radius, and eccentricity to put transiting planets detected around young stars in context. Bravo-Medina et al. (2021) and Mettler et al. (2020) used Exo-MerCat to quantify the amount of confirmed and candidate planets known to date.

Exo-MerCat is also used by the PLAnetary Transits and Oscillations of stars (PLATO) and the Large Interferometer For

Exoplanets (LIFE) science teams as source to their Input Catalogs.

PLATO is an adopted ESA mission planned for launch in 2026, focused on detecting and characterizing transiting exoplanets in the habitable zone of Sun-like stars (see e.g., Rauer et al., 2024, for details). In the PLATO Input Catalog (PIC), Exo-MerCat is used to identify all the stars known to host confirmed or candidate planets (including non-transiting ones) with the aim of including them for observations, even if their properties are not compliant with the criteria assumed for the PLATO scientific targets (see Montalto et al., 2021). Though the default Exo-MerCat-generated catalog does not contain information on the stellar host, it is essential to have a list of host star identifiers to query in the stellar catalogs. At the time of writing, the PIC developers have already upgraded the Exo-MerCat version to v2.0.0 and can benefit from the higher quality merge performed by Exo-MerCat v2.0.0 which allows to unambiguously identify each target with their original names and status identifiers in the exoplanet source catalogs. All of this allows them to gather additional information on the star from the exoplanet source catalogs and stellar catalogs through automated scripts and requiring little to no human supervision. The `duplicate_catalog_flag`, `duplicate_names`, `period_mismatch_flag` and `misnamed_duplicates_flag` columns allow them to make informed choices on potentially duplicated data or missed merges. In addition, Exo-MerCat was used to perform statistics on known and candidate planets in the LOPS2, i.e. the first Long Pointing PLATO field (see Nascimbeni et al. 2024, submitted).

LIFE is a concept for a space-based nulling interferometer, aiming at detecting and characterizing the thermal emission of terrestrial planets around F, G, K, and M stars in direct imaging (see e.g., Quanz et al., 2022, for details). The LIFE Target Database is being developed to provide a list of preliminary potential targets for the LIFE mission. Its goal is to merge information for stars, disks, and planets, that would be part of the LIFE Target List. Exo-MerCat is used to select all the stars known to host planets. It is especially useful to the team since the most challenging part, the comparison and merging of the data among various exoplanet catalog, has already been performed and homogenized. In future releases of the LIFE Target Database, Exo-MerCat will be used to prioritize targets for the LIFE observations, e.g., higher priority will be given to stars where interesting exoplanets are detected and lower priority to those that are known to host Hot Jupiters and where theory predicts no habitable terrestrial planets should be present.

The capability of Exo-MerCat of merging together different exoplanet catalogs, by homogenizing them, is also useful for webtools whose main aim is to search for exoplanetary system architectures. Among these we can cite the ASI-SSDC ExoAn3T (Exoplanetary Analysis and 3D Tool¹⁴) which, by querying remote archives, such as the NASA and EU catalogs, together with Exo-MerCat, is able to provide lists of exoplanetary systems based on user-defined planetary, stellar, or orbital

characteristics. ExoAn3T has already demonstrated its usefulness for the study of the dynamical history of a series of exosystems (e.g., Zinzi and Turrini (2017), Turrini et al. (2020), Turrini et al. (2023)) and can also provide hints in the field of astrobiology studies, when used to look for exosystems with architectures similar to the Solar System, or that host potentially habitable planets Zinzi et al. (2021).

The Exo-MerCat script can benefit a variety of astronomers: some might find the catalogs that are maintained and regularly updated to the TAP by the Exo-MerCat team useful; others might instead prefer to download the script, to execute it locally and make changes. In this case, quality control of the results is also passed down to the user.

3.3. Code support and future improvements

We released Exo-MerCat v2.0.0 on GitHub¹⁵ as open source, together with other useful modules (e.g., the Graphic User Interface, see Alei et al., 2020, for details). Users are encouraged to fork the repository and work with the code. Any update on the existing GitHub will be reviewed through the “merge request” process. This includes updates on the `replacements.ini` file, if further replacements are found when running the code.

The code runs with Python 3.8 and above and it is fully tested through unit tests. These are all expected to pass for a stable version of the script. In the case of an update of any input source, an Exo-MerCat run might still fail. The user can report the issue in the “Issues” panel.

The code has been fully documented. The main documentation can be found in the ReadTheDocs webpage¹⁶.

Both catalogs produced by Exo-MerCat v1.1.0 and Exo-MerCat v2.0.0 are reachable through a TAP service, hosted at IA2 Data Center¹⁷, in a similar way as described in (Alei et al., 2020). The TAP service is the same for both versions¹⁸ and the two versions are available as different resources. The catalog generated by Exo-MerCat v1.1.0 is stored as `exomercat.exomercat_legacy` and is no longer actively maintained, but users can still query the latest available update. Exo-MerCat v2.0.0 is executed monthly and two tables are uploaded on the TAP: `exomercat.exomercat`, which contains the result of the latest run, i.e. the most updated snapshot of the Exo-MerCat-generated catalog; and `exomercat.exomercat_global`, which collects all the entries that were updated on each run, without repeating identical entries from different runs. The structure of the two new tables is the same. Compared to the columns in the output `.csv` file of Exo-MerCat, both tables will feature an additional column, (`exomercat.version`) to log the version of the code that produced the result. This way, the user can easily retrieve the history of all modifications that a single target underwent in time. We plan also to store separately all the monthly runs of

¹⁴<https://tools.ssdsc.asi.it/exoplanet>

¹⁵<https://github.com/Exo-MerCat>

¹⁶<https://exo-mercat.readthedocs.io/en/latest/>

¹⁷<https://www.ia2.inaf.it/>

¹⁸<http://archives.ia2.inaf.it/vo/tap/projects>

Exo-MerCat v2.0.0. This will give us the possibility in the future to create and implement a service to allow users to download a snapshot of specific time in the past.

Future improvements on the code include: the addition of ancillary sources; the production of log files that are specific to the various input sources, which can be used to inform the maintainers of input sources of errors in the input catalogs in a more streamlined way; and the addition of further user cases and tutorials in the documentation. Exo-MerCat will be indexed in all major astronomic software databases such as the Exoplanet Modeling and Analysis Center (EMAC, Renaud et al., 2022) and the Astrophysics Source Code Library (ASCL, Nemiroff and Wallin, 1999; Allen et al., 2020) for a wider spread in the community.

4. Summary

In this work, we introduced the newest version of Exo-MerCat, a script that collects data from the most known exoplanet databases and homogenizes the sample in a VO-aware way. The most relevant upgrades that differentiate Exo-MerCat v2.0.0 from the previous versions of the code are summarized here:

- The TESS Project Candidates Table was added to the input sources, and the EPIC/K2 Planets and Candidates Table was promoted from look-up table to input source. The Exoplanet Orbit Database was removed from the input sources, since it has since been retired. The interface with NASA Exoplanet Archive, the Kepler Objects of Interest catalog, and the Exoplanet Encyclopaedia was also updated to comply with the changes in those databases. This upgrade increases the number of entries in the output catalog and the overall completeness of the parameter estimates in the sample.
- The main identifier is now searched in SIMBAD and TESS Input Catalog v8.2 for both name- and coordinate-based queries. In the case of coordinate-based queries, the query is performed on a much smaller tolerance radius (1 arcsec instead of 36 arcsec). This reduces the amount of mistakes in the attribution of the main identifier compared to the previous version.
- The grouping of the entries now takes into account the estimates of the period or semi-major axis. This ensures a more accurate merge that overcomes discrepancies in the planet letter.
- Exo-MerCat now produces a filtered version of the catalogs that removes all brown dwarf candidates with mass smaller than $20 M_{Jup}$, in addition to the complete catalog. Extra flags have been added to the final table and allow for custom filtering.
- The script produces more user-friendly log files, which can be used to force corrections to enable the correct merging of the entries. The user can easily handle these corrections

through ancillary files. This update adds more flexibility and user-friendliness to the script.

- The code has been restructured into modules and functions. This upgrade enables the inclusion of more input sources in the future.
- Exo-MerCat v2.0.0 has been released open-source and it is currently fully tested and documented.

Exo-MerCat has been used for various studies since the first publication. The catalog is also a source for the PLATO Input Catalog and the LIFE Target Database. The release of the new version of the code would be a valuable resource for other mission input catalogs, such as the Habitable Worlds Observatory Preliminary Input Catalog (HPIC, Tuchow et al., 2024). The open-source release of Exo-MerCat also allows for flexibility in the long-term maintenance of the code, which can be fully customized to the needs of every team.

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During the preparation of this work the authors used ChatGPT in order to produce code documentation, refactor code, and produce unit tests. After using this tool/service, the authors reviewed and edited the content as needed and take full responsibility for the content of the publication.

CRedit roles: E. Alei: Conceptualization, Data curation, Formal Analysis, Investigation, Methodology, Project administration, Software, Validation, Visualization, Writing - original draft, Writing - review and editing; S. Marinoni: Investigation, Methodology, Validation, Writing - original draft, Writing - review and editing; A. Bignamini: Data curation, Funding acquisition, Project administration, Resources, Software, Writing - original draft, Writing - review and editing; R. Claudi: Conceptualization, Funding acquisition, Project administration, Supervision, Writing - review and editing; M. Molinaro: Data curation, Resources, Software, Writing - review and editing; M. Vicinanza: Data curation, Software; S. Benatti: Writing - review and editing; I. Carleo: Writing - review and editing; A. Mandell: Writing - review and editing; F. Menti: Writing - review and editing; A. Zinzi: Writing - review and editing.

Appendix A. Column headers translation

In Table B.4 we list the previous and current naming convention for the catalog generated by Exo-MerCat. We indicate

with a star symbol (★) all of the columns that changed their meaning between the two versions of the script. We refer to the main text for more details on these changes.

Appendix B. Log for the October 25, 2024 run

As of October 25, 2024, the concatenated DataFrame contained 30034 entries, which were thus distributed: 9804 entries from EU, 7241 entries from TOI, 5785 entries from NASA, 5402 entries from OEC, 1800 entries from EPIC. Out of 15479 unique host star names, 2088 entries had as host star name one of the aliases in other entries and were set to the same value.

A total of 649 systems with an inconsistent value of “S-type” or “null” binary together with non-null values (641 occurrences) or only “S-type” and “null” occurrences (8 occurrences) could be fixed automatically. Of these, 370 systems (1085 single entries) showed coordinates that were outside the expected tolerance (`binary_coordinate_mismatch_flag=1`). This check identified also 19 complex systems (70 total entries with `binary_complex_system_flag=1`).

One entry (2MASS J0441+2301 A b) failed the coordinate check before the main identifier search for both right ascension and declination (with a separation of about 11 arcsec between the occurrences).

The main identifier search showed the following results: 28158 out of 30034 of the identifiers in the cumulative DataFrame were found by querying SIMBAD by host name or alias (94% sample completeness); 1524 more were found by querying TIC by host name or alias (98.9% sample completeness), 163 more were found via coordinates (99.4% sample completeness). Only 189 entries out of 30034 in the concatenated table could not be found in the stellar catalogs. For these, the identifiers were taken from the input catalogs.

The binary mismatch check on the main identifier allowed the fix of 19 more S-type systems, of which 13 had differences in the coordinates (`binary_coordinate_mismatch_flag=1`); 38 additional complex systems were found (`binary_complex_system_flag=1`). The main identifier suggested the presence of 28 potential binary stars that were not considered in the original sources. Six pairs of stars were found to have the same value of `host` but different values of `main_id`, either complex systems or stars are discovered to be gravitationally bound only after their discovery.

The planet letters were incoherent in 1094 occurrences. In 1020 of these, the planets in the cumulative catalog were labeled with their candidate string (“0d”) as well as their confirmed string (a lowercase letter between “b” and “j”). In these cases, Exo-MerCat forced the value of letter to be the one of the confirmed planet. The 74 remaining cases could not be automatically homogenized and were logged to be fixed manually.

The final table had 17745 entries, whose status is thus distributed: 7621 confirmed planets; 7632 candidate planets; 1488 false positives; 702 controversial planets; and two preliminary planets. When merging the entries to create the final catalog, 197 groups disagreed on the period or semi-major axis

(`period_mismatch_flag=1`) and resulted in 399 separate entries; 16 groups had a fallback merge. In total, these entries constituted about 2% of the final number of entries in the catalog generated by Exo-MerCat. A total of 62 entries had duplicate values in one or more input catalogs.

Table B.4: Column headers used by Exo-MerCat v1.0.0, Exo-MerCat v2.0.0, their meaning, and their type.

Exo-MerCat v1.0.0	Exo-MerCat v2.0.0	Meaning	Type
name	exo-mercat_name	Planet name chosen by Exo-MerCat	str
-	nasa_name	Planet name in the NASA Exoplanet Archive	str
-	toi_name	Planet name in the TESS Objects of Interest Table	str
-	epic_name	Planet name in the Kepler/K2 Objects of Interest	str
-	eu_name	Planet name in the Exoplanet Encyclopaedia	str
-	oec_name	Planet name in the Open Exoplanet Catalog	str
host	host	Host star name	str
letter	letter	Letter labeling the planet	str
main_id	main_id	★ Main identifier of the host star from SIMBAD or TIC catalogs	str
binary	binary	String labeling the binary host star, if any	str
ra_off	main_id_ra	★ J2000 right ascension in degrees from SIMBAD	float
dec_off	main_id_dec	★ J2000 declination in degrees from SIMBAD	float
mass	mass	Planet mass in Jovian masses	float
mass_max	mass_max	Positive error on the mass in Jovian masses	float
mass_min	mass_min	Negative error on the mass in Jovian masses	float
mass_url	mass_url	Bibcode of the reference paper for the mass value	str
msini	msini	Planet minimum mass in Jovian masses	float
msini_max	msini_max	Positive error on the minimum mass in Jovian masses	float
msini_min	msini_min	Negative error on the minimum mass in Jovian masses	float
msini_url	msini_url	Bibcode of the reference paper for the minimum mass value	str
bestmass	bestmass	Most precise value between mass and minimum mass in Jovian masses	float
bestmass_max	bestmass_max	Positive error on the best mass in Jovian masses	float
bestmass_min	bestmass_min	Negative error on the best mass in Jovian masses	float
bestmass_url	bestmass_url	Bibcode of the reference paper for the best mass value	str
mass_prov	bestmass_provenance	String labeling the origin of the best mass (mass or minimum mass)	str
p	p	Planet orbital period in days	float
p_max	p_max	Positive error on the period in days	float
p_min	p_min	Negative error on the period in days	float
p_url	p_url	Bibcode of the reference paper for the period value	str
r	r	Planet radius in Jovian radii	float
r_max	r_max	Positive error on the radius in Jovian radii	float
r_min	r_min	Negative error on the radius in Jovian radii	float
r_url	r_url	Bibcode of the reference paper for the radius value	str
a	a	Planet semi-major axis in au	float

a_max	a_max	Positive error on the semi-major axis in au	float
a_min	a_min	Negative error on the semi-major axis in au	float
a_url	a_url	Bibcode of the reference paper for the semi-major axis value	str
e	e	Eccentricity of the planet (scalar between 0 and 1)	float
e_max	e_max	Positive error on the eccentricity (scalar)	float
e_min	e_min	Negative error on the eccentricity (scalar)	float
e_url	e_url	Bibcode of the reference paper for the eccentricity value	str
i	i	Planet inclination in degrees.	float
i_max	i_max	Positive error on the inclination in degrees	float
i_min	i_min	Negative error on the minimum mass in degrees	float
i_url	i_url	Bibcode of the reference paper for the inclination value	str
discovery_method	discovery_method	Planet discovery method	str
Status	status	★ Planet status of the planet preferred by Exo-MerCat	str
Status_string	original_status_string	★ String listing the planet status in all available input sources	str
-	checked_status_string	★ String listing the planet status in all available input sources after the KOI check	str
confirmed	confirmed	★ Number of CONFIRMED status in checked_status_string	int
yod	discovery_year	Planet discovery year	int
alias	main_id_aliases	String listing all known aliases for the host star	str
-	main_id_provenance	Provenance of the main identifier	str
-	angular_separation_flag	Flag for non-null values of angular_separation	int
-	angular_separation	String listing the unique angular separations of the merged entries	str
catalog	catalog	String listing the catalogs in which the target appears	str
-	duplicate_catalog_flag	Flag for duplicates during merging	int
-	duplicate_names	String listing the planet name in all available input sources (only if duplicate_catalog_flag is not null)	str
MismatchFlagHost	binary_coordinate_mismatch_flag	★ Flag for possible coordinate mismatches in binaries	int
-	binary_complex_system_flag	★ Flag for complex binary systems	int
-	coordinate_mismatch_flag	Flag for possible coordinate mismatches	int
-	coordinate_mismatch	String labeling the coordinates for which a coordinate mismatch occurred, if any)	str
-	period_mismatch_flag	Flag for potential duplicates in final catalog because of mismatching p or a	int
-	fallback_merge_flag	Flag for potential duplicates in final catalog because no p or a are available (fallback merge case)	int
-	misnamed_duplicates_flag	Flag for potential duplicates in final catalog because of a mismatch in exo-mercat_name when p or a are the same	int
-	row_update	Date of the last update of the row.	str

References

- Akeson, R.L., Chen, X., Ciardi, D., Crane, M., Good, J., Harbut, M., Jackson, E., Kane, S.R., Laity, A.C., Leifer, S., Lynn, M., McElroy, D.L., Papin, M., Plavchan, P., Ramírez, S.V., Rey, R., von Braun, K., Wittman, M., Abajian, M., Ali, B., Beichman, C., Beekley, A., Berriman, G.B., Berukoff, S., Bryden, G., Chan, B., Groom, S., Lau, C., Payne, A.N., Regelson, M., Saucedo, M., Schmitz, M., Stauffer, J., Wyatt, P., Zhang, A., 2013. The NASA Exoplanet Archive: Data and Tools for Exoplanet Research. Publications of the Astronomical Society of the Pacific 125, 989–999. URL: <http://iopscience.iop.org/article/10.1086/672273/pdf><http://www.jstor.org/stable/info/10.1086/672273>, doi:10.1086/672273.
- Alei, E., Claudi, R., Bignamini, A., Molinaro, M., 2020. Exo-MerCat: A merged exoplanet catalog with Virtual Observatory connection. *Astronomy and Computing* 31, 100370. doi:10.1016/j.ascom.2020.100370, arXiv:2002.01834.
- Allen, A., Berriman, G.B., DuPrie, K., Mink, J., Nemiroff, R., Ryan, P.W., Schmidt, J., Shamir, L., Shortridge, K., Taylor, M., Teuben, P., Wallin, J., Warmels, R.H., 2020. The Astrophysics Source Code Library: What's New, What's Coming, in: Ballester, P., Ilsen, J., Solar, M., Shortridge, K. (Eds.), *Astronomical Data Analysis Software and Systems XXVII*, p. 731.
- Benatti, S., Damasso, M., Borsa, F., Locci, D., Pillitteri, I., Desidera, S., Maggior, A., Micela, G., Wolk, S., Claudi, R., Malavolta, L., Modirrousta-Galian, D., 2021. Constraints on the mass and on the atmospheric composition and evolution of the low-density young planet DS Tucanae A b. *Astronomy & Astrophysics* 650, A66. doi:10.1051/0004-6361/202140416, arXiv:2103.12922.
- Bravo-Medina, B., Strojnik, M., Mora-Núñez, A., Santiago-Hernández, H., 2021. Rotational-shearing-interferometer response for a star-planet system without star cancellation. *Applied Sciences* 11. URL: <https://www.mdpi.com/2076-3417/11/8/3322>, doi:10.3390/app11083322.
- Carleo, I., Desidera, S., Nardiello, D., Malavolta, L., Lanza, A.F., Livingston, J., Locci, D., Marzari, F., Messina, S., Turrini, D., Baratella, M., Borsa, F., D'Orazi, V., Nascimbeni, V., Pinamonti, M., Rainer, M., Alei, E., Bignamini, A., Gratton, R., Micela, G., Montalto, M., Sozzetti, A., Squicciarini, V., Affer, L., Benatti, S., Biazzo, K., Bonomo, A.S., Claudi, R., Cosentino, R., Covino, E., Damasso, M., Esposito, M., Fiorenzano, A., Frustagli, G., Giacobbe, P., Harutyunyan, A., Leto, G., Magazzù, A., Maggior, A., Mainella, G., Maldonado, J., Mallonn, M., Mancini, L., Molinari, E., Molinaro, M., Pagano, I., Pedani, M., Piotto, G., Poretti, E., Redfield, S., Scandariato, G., 2021. The GAPS Programme at TNG. XXVIII. A pair of hot-Neptunes orbiting the young star TOI-942. *Astronomy & Astrophysics* 645, A71. doi:10.1051/0004-6361/202039042, arXiv:2011.13795.
- Desidera, S., Damasso, M., Gratton, R., Benatti, S., Nardiello, D., D'Orazi, V., Lanza, A.F., Locci, D., Marzari, F., Mesa, D., Messina, S., Pillitteri, I., Sozzetti, A., Girard, J., Maggio, A., Micela, G., Malavolta, L., Nascimbeni, V., Pinamonti, M., Squicciarini, V., Alcalá, J., Biazzo, K., Bohn, A., Bonavita, M., Brooks, K., Chauvin, G., Covino, E., Delorme, P., Hagelberg, J., Janson, M., Lagrange, A.M., Lazzoni, C., 2023. TOI-179: A young system with a transiting compact Neptune-mass planet and a low-mass companion in outer orbit. *Astronomy & Astrophysics* 675, A158. doi:10.1051/0004-6361/202244611, arXiv:2210.07933.
- Gaia Collaboration, Brown, A.G.A., Vallenari, A., Prusti, T., de Bruijne, J.H.J., Babusiaux, C., Bailer-Jones, C.A.L., Biermann, M., Evans, D.W., Eyer, L., et al., 2018. Gaia Data Release 2. Summary of the contents and survey properties. *A&A* 616, A1. doi:10.1051/0004-6361/201833051, arXiv:1804.09365.
- Gaia Collaboration, Prusti, T., de Bruijne, J.H.J., Brown, A.G.A., Vallenari, A., Babusiaux, C., Bailer-Jones, C.A.L., Bastian, U., Biermann, M., Evans, D.W., et al., 2016. The Gaia mission. *A&A* 595, A1. doi:10.1051/0004-6361/201629272, arXiv:1609.04153.
- Huber, D., Bryson, S.T., et al., 2017. VizieR Online Data Catalog: K2 Ecliptic Plane Input Catalog (EPIC) (Huber+, 2017). *VizieR Online Data Catalog*, IV/34.
- Kepler Mission Team, 2009. *VizieR Online Data Catalog: Kepler Input Catalog* (Kepler Mission Team, 2009). *VizieR Online Data Catalog*, V/133.
- Mayor, M., Queloz, D., 1995. A Jupiter-mass companion to a solar-type star. *Nature* 378, 355–359. doi:10.1038/378355a0.
- Mettler, J.N., Quanz, S.P., Helled, R., 2020. Earth as an Exoplanet. I. Time Variable Thermal Emission Using Spatially Resolved Moderate Imaging Spectroradiometer Data. *The Astronomical Journal* 160, 246. doi:10.3847/1538-3881/abb15, arXiv:2010.02589.
- Montalto, M., Piotto, G., Marrese, P.M., Nascimbeni, V., Prisinzano, L., Granata, V., Marinoni, S., Desidera, S., Ortolani, S., Aerts, C., Alei, E., Altavilla, G., Benatti, S., Börner, A., Cabrera, J., Claudi, R., Deleuil, M., Fabrizio, M., Gizon, L., Goupil, M.J., Heras, A.M., Magrin, D., Malavolta, L., Mas-Hesse, J.M., Pagano, I., Paproth, C., Pertenais, M., Pollacco, D., Ragazzoni, R., Ramsay, G., Rauer, H., Udry, S., 2021. The all-sky PLATO input catalogue. *A&A* 653, A98. doi:10.1051/0004-6361/202140717, arXiv:2108.13712.
- Nemiroff, R.J., Wallin, J.F., 1999. The Astrophysics Source Code Library: <http://www.ascl.net/>, in: *American Astronomical Society Meeting Abstracts* #194, p. 44.08.
- Paegert, M., Stassun, K.G., Collins, K.A., Pepper, J., Torres, G., Jenkins, J., Twicken, J.D., Latham, D.W., 2021. TESS Input Catalog versions 8.1 and 8.2: Phantoms in the 8.0 Catalog and How to Handle Them. arXiv e-prints, arXiv:2108.04778doi:10.48550/arXiv.2108.04778, arXiv:2108.04778.
- Quanz, S.P., Ottiger, M., Fontanet, E., Kammerer, J., Menti, F., Dannert, F., Gheorghe, A., Absil, O., Airapetian, V.S., Alei, E., Allart, R., Angerhausen, D., Blumenthal, S., Buchhave, L.A., Cabrera, J., Carrión-González, Ó., Chauvin, G., Danchi, W.C., Dandmont, C., Defrére, D., Dorn, C., Ehrenreich, D., Ertel, S., Fridlund, M., García Muñoz, A., Gascón, C., Girard, J.H., Glauser, A., Grenfell, J.L., Guidi, G., Hagelberg, J., Helled, R., Ireland, M.J., Janson, M., Kopparapu, R.K., Korth, J., Kozakis, T., Kraus, S., Léger, A., Leedjävär, L., Lichtenberg, T., Lillo-Box, J., Linz, H., Liseau, R., Loicq, J., Mahendra, V., Malbet, F., Mathew, J., Mennesson, B., Meyer, M.R., Mishra, L., Molaverdikhani, K., Noack, L., Oza, A.V., Pallé, E., Parviainen, H., Quirrenbach, A., Rauer, H., Ribas, I., Rice, M., Romagnolo, A., Rugheimer, S., Schwieterman, E.W., Serabyn, E., Sharma, S., Stassun, K.G., Szulágyi, J., Wang, H.S., Wunderlich, F., Wyatt, M.C., LIFE Collaboration, 2022. Large Interferometer For Exoplanets (LIFE). I. Improved exoplanet detection yield estimates for a large mid-infrared space-interferometer mission. *Astronomy & Astrophysics* 664, A21. doi:10.1051/0004-6361/202140366, arXiv:2101.07500.
- Rauer, H., Aerts, C., Cabrera, J., Deleuil, M., Erikson, A., Gizon, L., Goupil, M., Heras, A., Lorenzo-Alvarez, J., Marliani, F., Martin-Garcia, C., Mas-Hesse, J.M., O'Rourke, L., Osborn, H., Pagano, I., Piotto, G., Pollacco, D., Ragazzoni, R., Ramsay, G., Udry, S., Apourchaux, T., Benz, W., Brander, A., Güdel, M., Janot-Pacheco, E., Kabath, P., Kjeldsen, H., Min, M., Santos, N., Smith, A., Suarez, J.C., Werner, S.C., Aboudan, A., Abreu, M., Acuña, L., Adams, M., Adibekyan, V., Affer, L., Agneray, F., Agnor, C., Aguirre Børsen-Koch, V., Ahmed, S., Aigrain, S., Al-Bahlawan, A., Alcacera Gil, M.d.I.A., Alei, E., Alencar, S., Alexander, R., Alfonso-Garzón, J., Alibert, Y., Allende Prieto, C., Almeida, L., Alonso Sobrino, R., Altavilla, G., Althaus, C., Alonso Alvarez Trujillo, L., Amarsi, A., Ammler-von Eiff, M., Amôres, E., Andrade, L., Antoniadis-Karnavas, A., António, C., Aparicio del Moral, B., Appolloni, M., Arena, C., Armstrong, D., Aroca Aliaga, J., Asplund, M., Audenaert, J., Auricchio, N., Avelino, P., Baeke, A., Baillié, K., Balado, A., Balestra, A., Ball, W., Ballans, H., Ballot, J., Barban, C., Barbary, G., Barbieri, M., Barceló Forteza, S., Barker, A., Barklem, P., Barnes, S., Barrado Navascues, D., Barragan, O., Baruteau, C., Basu, S., Baudin, F., Baumeister, P., Bayliss, D., Bazot, M., Beck, P.G., Bedding, T., Belkacem, K., Bellinger, E., Benatti, S., Benomar, O., Bérard, D., Berge-mann, M., Bergomi, M., Bernardo, P., Biazzo, K., Bignamini, A., Bigot, L., Billot, N., Binet, M., Biondi, D., Biondi, F., Birch, A.C., Bitsch, B., Bluhm Ceballos, P.V., Bódi, A., Bognár, Z., Boisse, I., Bolmont, E., Bonanno, A., Bonavita, M., Bonfanti, A., Bonfils, X., Bonito, R., Bonomo, A.S., Börner, A., Boro Saikia, S., Borroguero Martín, E., Borsa, F., Borsato, L., Bossini, D., Bouchy, F., Boué, G., Bouffleur, R., Boumier, P., Bourrier, V., Bowman, D.M., Bozzo, E., Bradley, L., Bray, J., Bressan, A., Breton, S., Brienza, D., Brito, A., Brogi, M., Brown, B., Brown, D., Brun, A.S., Bruno, G., Bruns, M., Buchhave, L.A., Bugnet, L., Buldgen, G., Burgess, P., Busatta, A., Busso, G., Buzasi, D., Caballero, J.A., Cabral, A., Calderone, F., Cameron, R., Cameron, A., Campante, T., Canto Martins, B.L., Cara, C., Carone, L., Carrasco, J.M., Casagrande, L., Casewell, S.L., Cassisi, S., Castellani, M., Castro, M., Catala, C., Catalán Fernández, I., Catelan, M., Cegla, H., Cerruti, C., Cessa, V., Chadid, M., Chaplin, W., Charpinet, S., Chiappini, C., Chiarucci, S., Chiavassa, A., Chinellato, S., Chirulli, G., Christensen-Dalsgaard, J., Church, R., Claret, A., Clarke, C., Claudi, R., Clermont, L., Coelho, H., Coelho, J., Cogato, F., Colomé, J., Condamine,

- M., Conseil, S., Corbard, T., Correia, A.C.M., Corsaro, E., Cosentino, R., Costes, J., Cottinelli, A., Covone, G., Creevey, O.L., Crida, A., Csizmadia, S., Cunha, M., Curry, P., da Costa, J., da Silva, F., Dalal, S., Damasso, M., Damiani, C., Damiani, F., Liduina das Chagas, M., Davies, M., Davies, G., Davies, B., Davison, G., de Almeida, L., de Angeli, F., Cabral de Barros, S.C., de Castro Leão, I., Brito de Freitas, D., de Freitas, M.C., De Martino, D., Renan de Medeiros, J., de Paula, L.A., de Plaa, J., De Ridder, J., Deal, M., Decin, L., Deeg, H., Degl'Innocenti, S., Deheuvels, S., del Burgo, C., Del Sordo, F., Delgado-Mena, E., Demangeon, O., Denk, T., Derekas, A., Desidera, S., Dexet, M., Di Criscienzo, M., Di Giorgio, A.M., Di Mauro, M.P., Diaz Rial, F.J., Díaz-García, J.J., Dima, M., Dinuzzi, G., Dionatos, O., Distefano, E., do Nascimento, Jose-Dias, J., Domingo, A., D'Orazi, V., Dorn, C., Doyle, L., Duarte, E., Ducellier, F., Dumaye, L., Dumusque, X., Dupret, M.A., Eggenberger, P., Ehrenreich, D., Eiglmüller, P., Eising, J., Emilio, M., Eriksson, K., Ermocida, M., Isidoro Escate Giribaldi, R., Eschen, Y., Estrela, I., Evans, D.W., Fabbian, D., Fabrizio, M., Faria, J.P., Farina, M., Farinato, J., Feliz, D., Feltzing, S., Fenouillet, T., Ferrari, L., Ferraz-Mello, S., Fialho, F., Fienga, A., Figueira, P., Fiori, L., Flacomio, E., Focardi, M., Foley, S., Fontignie, J., Ford, D., Fornazier, K., Forveille, T., Fossati, L., de Marca Franca, R., da Silva, L.F., Frasca, A., Fridlund, M., Furlan, M., Gabler, S.M., Gaido, M., Gallagher, A., Galli, E., García, R.A., García Hernández, A., García Munoz, A., García-Vázquez, H., Garrido Haba, R., Gaulme, P., Gauthier, N., Gehan, C., Gent, M., Georgieva, I., Ghigo, M., Giana, E., Gill, S., Girardi, L., Giuliani Winter, S., Giusi, G., Gomes da Silva, J., Gómez Zazo, L.J., Gomez-Lopez, J.M., Isai González Hernández, J., Gonzalez Murillo, K., Gorius, N., Gouel, P.V., Goulty, D., Granata, V., Grenfell, J.L., Griebbach, D., Grolleau, E., Grouffal, S., Grziwa, S., Guarcello, M.G., Gueguen, L., Guenther, E.W., Guilhem, T., Guillerot, L., Guiot, P., Guterman, P., Gutiérrez, A., Gutiérrez-Canales, F., Hagelberg, J., Haldemann, J., Hall, C., Handberg, R., Harrison, I., Harrison, D.L., Hasiba, J., Haswell, C.A., Hatalova, P., Hatzes, A., Haywood, R., Hébrard, G., Heckes, F., Heiter, U., Hekker, S., Heller, R., Helling, C., Helminiak, K., Hemsley, S., Heng, K., Hermans, A., Hermes, J., Hidalgo Torres, N., Hinkel, N., Hobbs, D., Hodgkin, S., Hofmann, K., Hojjatpanah, S., Houdek, G., Huber, D., Huesler, J., Hui-Bon-Hoa, A., Huygen, R., Huynh, D.D., Iro, N., Irwin, J., Irwin, M., Izidoro, A., Jacquinod, S., Emborg Jannsen, N., Janson, M., Jeszenszky, H., Jiang, C., José Jimenez Mancebo, A., Jofre, P., Johansen, A., Johnston, C., Jones, G., Kallinger, T., Kálmán, S., Kanitz, T., Karjalainen, M., Karjalainen, R., Karoff, C., Kawaler, S., Kawata, D., Keereman, A., Keiderling, D., Kennedy, T., Kenworthy, M., Kerschbaum, F., Kidger, M., Kiefer, F., Kintziger, C., Kislyakova, K., Kiss, L., Klagyivik, P., Klahr, H., Klevas, J., Kochukhov, O., Köhler, U., Kolb, U., Koncz, A., Korth, J., Kostogryz, N., Kovács, G., Kovács, J., Kozhura, O., Krivova, N., Kučinskas, A., Kuhlemann, I., Kupka, F., Laauwen, W., Labiano, A., Lagarde, N., Laget, P., Laky, G., Lam, K.W.F., Lambrechts, M., Lammer, H., Lanza, A.F., Lanzafame, A., Lares Martiz, M., Laskar, J., Latter, H., Lavanant, T., Lawrenson, A., Lazzoni, C., Lebre, A., Lebreton, Y., Lecavelier des Etangs, A., Leinhardt, Z., Leleu, A., Lendl, M., Leto, G., Levillain, Y., Libert, A.S., Lichtenberg, T., Ligi, R., Lignieres, F., Lillo-Box, J., Linsky, J., Seige Liu, J., Loidolt, D., Longval, Y., Lopes, I., Lorenzani, A., Ludwig, H.G., Lund, M., Sloth Lundkvist, M., Luri, X., Maceroni, C., Madden, S., Madhusudhan, N., Maggio, A., Magliano, C., Magrin, D., Mahy, L., Maibaum, O., Malac-Allain, L., Malapert, J.C., Malavolta, L., Maldonado, J., Mamonova, E., Manchon, L., Mann, A., Mantovan, G., Marafatto, L., Marconi, M., Mardling, R., Marigo, P., Marinoni, S., Marques, É., Marques, J.P., Marrese, P.M., Marshall, D., Martínez Perales, S., Mary, D., Marzari, F., Masana, E., Mascher, A., Mathis, S., Mathur, S., Mattiuci Figueiredo, A.C., Maxted, P.F.L., Mazeh, T., Mazevet, S., Mazzei, F., McCormac, J., McMillan, P., Menou, L., Merle, T., Meru, F., Mesa, D., Messina, S., Mészáros, S., Meunier, N., Meunier, J.C., Micela, G., Michaelis, H., Michel, E., Michielsen, M., Michtchenko, T., Miglio, A., Miguel, Y., Milligan, D., Mirouh, G., Mitchell, M.A., Moedas, N., Molendini, F., Molnár, L., Mombarg, J., Montalbán, J., Montalto, M., Monteiro, M.J.P.F.G., Morales, J.C., Morales-Calderon, M., Morbidelli, A., Mordasini, C., Moreau, C., Morel, T., Morello, G., Morin, J., Mortier, A., Mosser, B., Mourad, D., Mousis, O., Moutou, C., Mowlavi, N., Moya, A., Muehlmann, P., Muirhead, P., Munari, M., Musella, I., Mustill, A.J., Nardetto, N., Nardiello, D., Narita, N., Nascimbeni, V., Nash, A., Neiner, C., Nelson, R.P., Nettelmann, N., Nicolini, G., Nielsen, M., Niemi, S.M., Noack, L., Noels-Grotsch, A., Noll, A., Norzaman, A., Norton, A.J., Nsamba, B., Ofir, A., Ogilvie, G., Olander, T., Olivetto, C., Olofsson, G., Ong, J., Ortolani, S., Oshagh, M., Ottacher, H., Ottensamer, R., Ouazzani, R.M., Paardekooper, S.J., Pace, E., Pajas, M., Palacios, A., Palandri, G., Palle, E., Paproth, C., Parro, V., Parviainen, H., Granado, J.P., Passegger, V.M., Pastor-Morales, C., Pätzold, M., Gade Pedersen, M., Pena Hidalgo, D., Pepe, F., Pereira, F., Persson, C.M., Pertenais, M., Peter, G., Petit, A.C., Petit, P., Pezzuto, S., Pichierrri, G., Pietrinferni, A., Pinheiro, F., Pinsonneault, M., Plachy, E., Plasson, P., Plez, B., Poppenhaeger, K., Poretti, E., Portaluri, E., Portell, J., Frederico Porto de Mello, G., Poyatos, J., Pozuelos, F.J., Prada Moroni, P.G., Pricopi, D., Prisinzano, L., Quade, M., Quirrenbach, I., Rabanal Reina, J.A., Rabello Soares, M.C., Raimondo, G., Rainer, M., Ramón Rodón, J., Ramón-Ballesta, A., Ramos Zapata, G., Rätz, S., Rauterberg, C., Redman, B., Redmer, R., Reese, D., Regibo, S., Reiners, A., Reinhold, T., Renie, C., Ribas, I., Ribeiro, S., Pereira Ricciardi, T., Rice, K., Richard, O., Riello, M., Rietord, M., Ripepi, V., Rixon, G., Rockstein, S., Rodríguez, M.T.R., Rodríguez Díaz, L.F., Rodríguez García, J.P., Rodríguez-Gómez, J., Roehly, Y., Roig, F., Rojas-Ayala, B., Rolf, T., Lysgaard Rørsted, J., Rosado, H., Rosotti, G., Roth, O., Roth, M., Rousseau, A., Roxburgh, I., Roy, F., Royer, P., Ruane, K., Rufini Mastropasqua, S., Ruiz de Galarreta, C., Russi, A., Saar, S., Saillenfest, M., Salaris, M., Salmon, S., Saltas, I., Samadi, R., Samadi, A., Samra, D., Sanches da Silva, T., Andrés Sánchez Carrasco, M., Santerne, A., Santoli, F., Santos, Á.R.G., Sanz Mesa, R., Sarro, L.M., Scandariato, G., Schäfer, M., Schlafly, E., Schmäder, F.X., Schneider, J., Schou, J., Schunker, H., Jörg Schwarzkopf, G., Serenelli, A., Seynaeve, D., Shaw, Y., Shapiro, A., Shipman, R., Sicilia, D., Sierra Sanmartín, M.A., Sigot, A., Silliman, K., Silvotti, R., Simon, A.E., Simoyama Napoli, R., Skarka, M., Smalley, B., Smiljanic, R., Smit, S., Smith, A., Smith, L., Snellen, I., Sódor, Á., Sohl, F., Solanki, S.K., Sortino, F., Sousa, S., Southworth, J., Souto, D., Sozzetti, A., Stamatellos, D., Stassun, K., Steller, M., Stello, D., Stelzer, B., Stiebel, U., Stokholm, A., Storeymo, T., Strassmeier, K., Ström, P.A., Strugarek, A., Sulis, S., Švanda, M., Szabados, L., Szabó, R., Szabó, G.M., Szuszkiewicz, E., Talens, G.J., Teti, D., Theisen, T., Thévenin, F., Thoul, A., Tiphene, D., Titz-Weider, R., Tkachenko, A., Tomecki, D., Tonfát, J., Tosi, N., Trampedach, R., Traven, G., Triaud, A., Trønnes, R., Tsantaki, M., Tschentscher, M., Turin, A., Tvaruzka, A., Ulmer, B., Ulmer-Moll, S., Ulusoy, C., Umbriaco, G., Valencia, D., Valentini, M., Valio, A., Valverde Guisjarro, Á.L., Van Eylen, V., Van Grootel, V., van Kempen, T.A., Van Reeth, T., Van Zelst, I., Vandenbussche, B., Vasilou, K., Vasilyev, V., Vaz de Mascarenhas, D., Vazan, A., Vela Nunez, M., Nunes Velloso, E., Ventura, R., Ventura, P., Venturini, J., Trallero, I.V., Veras, D., Verdugo, E., Verma, K., Vibert, D., Vicanek Martinez, T., Vida, K., Vigan, A., Villacorta, A., Villaver, E., Villaverde Aparicio, M., Viotto, V., Vorobyov, E., Vorontsov, S., Wagner, F.W., Walloschek, T., Walton, N., Walton, D., Wang, H., Waters, R., Watson, C., Wedemeyer, S., Weeks, A., Weingrill, J., Weiss, A., Wendler, B., West, R., Westerdorff, K., Westphal, P.A., Wheatley, P., White, T., Whittaker, A., Wickhusen, K., Wilson, T., Windsor, J., Winter, O., Lykke Winther, M., Winton, A., Witteck, U., Witzke, V., Woitke, P., Wolter, D., Wuchterl, G., Wyatt, M., Yang, D., Yu, J., Zanmar Sanchez, R., Rosa Zapatero Osorio, M., Zechmeister, M., Zhou, Y., Ziemke, C., Zwintz, K., 2024. The PLATO Mission. arXiv e-prints, arXiv:2406.05447doi:10.48550/arXiv.2406.05447, arXiv:2406.05447.
- Rein, H., 2012. A proposal for community driven and decentralized astronomical databases and the Open Exoplanet Catalogue. arXiv e-prints , arXiv:1211.7121arXiv:1211.7121.
- Renaud, J.P., Lopez, E., Brande, J., Cruz-Arce, C.E., Kelahan, C., Susemihl, N., Cristy, D., Hostetter, C., Moore, M.D., Patel, A., Mandell, A.M., 2022. The Exoplanet Modeling and Analysis Center at NASA Goddard. Research Notes of the American Astronomical Society 6, 185. doi:10.3847/2515-5172/ac9060, arXiv:2209.04005.
- Schneider, J., Dedieu, C., Le Sidaner, P., Savalle, R., Zolotukhin, I., 2011. Defining and cataloging exoplanets: the exoplanet.eu database. Astronomy & Astrophysics 532, A79. URL: <http://arxiv.org/abs/1106.0586>, doi:10.1051/0004-6361/201116713.
- Tuchow, N.W., Stark, C.C., Mamajek, E., 2024. Hpic: The habitable worlds observatory preliminary input catalog. The Astronomical Journal 167, 139. URL: <https://dx.doi.org/10.3847/1538-3881/ad25ec>, doi:10.3847/1538-3881/ad25ec.
- Turrini, D., Marzari, F., Polychroni, D., Claudi, R., Desidera, S., Mesa, D., Pinamonti, M., Sozzetti, A., Suárez Mascareño, A., Damasso, M., Benatti, S., Malavolta, L., Micela, G., Zanzi, A., Béjar, V.J.S., Biazzo, K., Bignamini, A., Bonavita, M., Borsari, F., del Burgo, C., Chauvin, G., Delorme, P., González Hernández, J.I., Gratton, R., Hagelberg, J., Janson, M., Lan-

- glois, M., Lanza, A.F., Lazzoni, C., Lodieu, N., Maggio, A., Mancini, L., Molinari, E., Molinaro, M., Murgas, F., Nardiello, D., 2023. The GAPS programme at TNG. XLVIII. The unusual formation history of V1298 Tau. *Astronomy & Astrophysics* 679. doi:10.1051/0004-6361/202244752, arXiv:2307.08653.
- Turrini, D., Zinzi, A., Belinchon, J.A., 2020. Normalized angular momentum deficit: a tool for comparing the violence of the dynamical histories of planetary systems. *Astronomy & Astrophysics* 636. doi:10.1051/0004-6361/201936301, arXiv:2003.05366.
- Wright, J.T., et al., 2011. The exoplanet orbit database. *Publications of the Astronomical Society of the Pacific* 123, 412. doi:10.1086/659427, arXiv:1012.5676.
- Zinzi, A., Turrini, D., 2017. Anti-correlation between multiplicity and orbital properties in exoplanetary systems as a possible record of their dynamical histories. *Astronomy & Astrophysics* 605. doi:10.1051/0004-6361/201731595, arXiv:1709.00003.
- Zinzi, A., Turrini, D., Alei, E., Verrecchia, F., 2021. Exoplan3t: a new way of exploring large exoplanetary databases and its applications to astrobiology. *Memorie della Società Astronomica Italiana* 92.